

HIGHLY ANISOTROPIC TEMPERATURE BALANCE EQUATION AND ITS ASYMPTOTIC-PRESERVING RESOLUTION.

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ABSTRACT. This paper deals with the numerical study of a nonlinear, strongly anisotropic heat equation. The use of standard schemes in this situation leads to poor results, due to the high anisotropy. An Asymptotic-Preserving method is introduced in this paper, which is second-order accurate in both, temporal and spacial variables. The discretization in time is done using an L-stable Runge-Kutta scheme. The convergence of the method is shown to be independent of the anisotropy parameter $0 < \varepsilon < 1$, and this for fixed coarse Cartesian grids and for variable anisotropy directions. The context of this work are magnetically confined fusion plasmas.

1. INTRODUCTION

Magnetically confined plasmas are characterized by highly anisotropic properties induced by the applied strong magnetic field. Indeed, the charged particles constituting the plasma move rapidly around the magnetic field lines, their transverse motion away from the field lines being constrained by the Lorentz force. In contrast, their motion along the field lines is relatively unconstrained, so that rather rapid dynamics along the magnetic fields occurs. This results in an extremely large ratio of the parallel to the transverse thermal conductivities, as well as of other parameters characterizing the plasma evolution.

A prototype simplified model for the heat diffusion in a magnetically confined plasma can be expressed by the following nonlinear, degenerate parabolic equation

$$\partial_t u - \nabla_{||} \cdot (\kappa_{||}(u) \nabla_{||} u) - \nabla_{\perp} \cdot (\kappa_{\perp} \nabla_{\perp} u) = 0, \quad (1.1)$$

where the subscripts $||$ (resp. \perp) refer to the direction parallel (resp. perpendicular) to the magnetic field lines and u designates the temperature. In writing out the equation above we have ignored some important physical phenomena coming from convection and turbulence. Nevertheless, our equation contains some important features inherited from the full model that lead to substantial difficulties in the numerical treatment of both the full model and our simplified one. The diffusion in the direction perpendicular to the magnetic field lines is usually dominated by the anomalous transport and the corresponding coefficient κ_{\perp} can be taken temperature independent. On the other hand, the coefficient describing the diffusion in the direction parallel to the magnetic field lines, $\kappa_{||}$, is normally much larger and strongly

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temperature dependent. It can be described by the Spitzer-Härm law $\kappa_{||}(u) = \kappa_0 u^{5/2}$ [24]. Moreover, plasma temperatures are extremely high, so that this diffusion coefficient can become very big. Passing to non-dimensional variables, we shall write therefore the law for $\kappa_{||}$ as

$$\kappa_{||}(u) = \frac{1}{\varepsilon} u^{5/2},$$

where ε is a small parameter, $0 < \varepsilon \ll 1$. An accurate resolution of the parallel and perpendicular diffusion processes plays a crucial role in understanding of the plasma dynamics and the energy transport phenomena. It is therefore very important to develop and to study efficient numerical schemes to solve problem (1.1). It is also desirable to have a scheme that works robustly for all values of ε from $\varepsilon \ll 1$ to $\varepsilon \sim O(1)$ since this parameter enters the equation in combination with a non-linear term so that the effective value of the diffusion coefficient can vary strongly over the computation domain following the variations in u . This is the primary motivation of the present work.

Anisotropic, nonlinear diffusion equations of the type (1.1) arise in several other fields of application and a lot of efforts were made to construct efficient numerical methods for this challenging problem. To mention some examples, such non-linear evolution equations of parabolic type occur in the description of isentropic gas flows through a porous media [2] or in the description of transport phenomena in heterogeneous geologic formations, such as fractured rock systems [5], which are of fundamental interest for petroleum or groundwater engineering. In addition, these equations appear also in image processing, related to the elimination of noise and small-scale details from an image [3, 18, 23] or in the description of the anisotropic water diffusion in tissues of the nervous system [4].

From a numerical point of view, problems of the type (1.1) are very challenging, as one deals with singularly perturbed problems, the model changing its type in the limit $\varepsilon \rightarrow 0$. Standard schemes suffer from the presence of very ill conditioned matrices (typically with a condition number of order $1/(\varepsilon h^2)$ where h is the discretization step in space). Solving an equation with such a matrix on a computer accumulates the rounding errors and may lead to completely wrong results. Note that this drawback cannot be overcome by a mesh refinement since it results only in worsening the condition numbers of the matrices in the discretized problem.

Several methods were investigated in literature to cope with this type of anisotropic problems, using for example high order finite element schemes [10], preconditioned conjugate gradient methods in a mixed spectral/finite difference scheme [15] or introducing an artificial “sound” method, to represent the fast thermal equilibrium along the field lines [17]. All these methods however are rather involved and moreover their range of application is limited, as they are efficient only until a threshold value for ε , and cannot thus recover the limit regime $\varepsilon \rightarrow 0$. Another class of employed numerical methods are hybrid strategies, which consist in coupling different numerical schemes valid in different regions

of the domain. For example in this case, one can couple the resolution of the singular perturbation problem there where $\varepsilon \sim \mathcal{O}(1)$ with the resolution of a limit problem for $\varepsilon \ll 1$. These methods suffer however from the fact, that the coupling conditions between the two models are hard to establish and the interface between the two regions difficult to localize.

The objective of the present paper is to introduce an efficient numerical scheme based on the Asymptotic-Preserving methodology, which allows for an accurate resolution of the singularly perturbed problem, uniformly in ε , with little additional computational cost, and using a grid which is not necessarily aligned with the magnetic field, so that one can exploit simple Cartesian grids, for example. Initially, AP-techniques were introduced in [12], to deal with singularly perturbed kinetic models. The key idea is to reformulate the singularly perturbed problem into an equivalent problem, which is however well-posed if we set $\varepsilon = 0$ there. The reformulation of the here proposed method is based, similarly as in [16], on introducing a new auxiliary variable, as proposed earlier in an elliptic framework in [7], and replacing the terms of the equation multiplied by $1/\varepsilon$ are by the new terms with an $O(1)$ factor. From a numerical point of view, this procedure means transforming the problem of condition number $\sim 1/\varepsilon$, into a well-conditioned problem, which switches automatically from the singularly perturbed problem to the limit problem, as $\varepsilon \rightarrow 0$.

The difference between the method presented in [16] lies first in the treatment of the non-linearity. Instead of fixed point iterations used to approach the nonlinearity $(u^{n+1})^{5/2}$, we choose to implement a much simpler linear extrapolation method (see Section 3.2 for details). Moreover, we develop here a robust asymptotic-preserving scheme of second order in time, which has no analogue in the existing literature, to the best of our knowledge. Finally, this paper contains also a detailed mathematical study of the problem.

The paper is organized as follows: Section 2 contains a description of the problem completed by a mathematical study. In Section 3, we present the numerical method based on an asymptotic preserving space discretization and develop three different time-discretizations: implicit Euler, Crank-Nicolson and L-stable Runge-Kutta methods. Finally, in Section 4 we present some numerical results, focusing on the AP-property of the schemes.

2. DESCRIPTION OF THE PROBLEM AND MATHEMATICAL STUDY

We consider a two or three dimensional anisotropic, nonlinear heat problem, given on a sufficiently smooth, bounded domain $\Omega \subset \mathbb{R}^d$, $d = 2, 3$ with boundary Γ . The direction of the anisotropy is defined by the time-independent vector field $b \in (C^\infty(\Omega))^d$, satisfying $|b(x)| = 1$ for all $x \in \Omega$.

Given this vector field b , one can decompose now vectors $v \in \mathbb{R}^d$, gradients $\nabla\phi$, with $\phi(x)$ a scalar function, and divergences $\nabla \cdot v$, with $v(x)$ a vector field, into a part parallel to the

anisotropy direction and a part perpendicular to it. These parts are defined as follows:

$$\begin{aligned} v_{\parallel} &:= (v \cdot b)b, & v_{\perp} &:= (Id - b \otimes b)v, & \text{such that } v &= v_{\parallel} + v_{\perp}, \\ \nabla_{\parallel}\phi &:= (b \cdot \nabla\phi)b, & \nabla_{\perp}\phi &:= (Id - b \otimes b)\nabla\phi, & \text{such that } \nabla\phi &= \nabla_{\parallel}\phi + \nabla_{\perp}\phi, \\ \nabla_{\parallel} \cdot v &:= \nabla \cdot v_{\parallel}, & \nabla_{\perp} \cdot v &:= \nabla \cdot v_{\perp}, & \text{such that } \nabla \cdot v &= \nabla_{\parallel} \cdot v + \nabla_{\perp} \cdot v, \end{aligned} \quad (2.2)$$

where we denoted by \otimes the vector tensor product.

The boundary Γ can be decomposed into three components following the sign of the intersection with b :

$$\Gamma_{\parallel} := \{x \in \Gamma / b(x) \cdot n(x) = 0\},$$

$$\Gamma_{in} := \{x \in \Gamma / b(x) \cdot n(x) < 0\}, \quad \Gamma_{out} := \{x \in \Gamma / b(x) \cdot n(x) > 0\},$$

and $\Gamma_{\perp} = \Gamma_{in} \cup \Gamma_{out}$. The vector n is here the unit outward normal on Γ .

With these notations we can now introduce the mathematical problem, we are interested to study. We are searching for the particle (ions or electrons) temperature $u(t, x)$, solution of the evolution equation

$$(P) \quad \left\{ \begin{array}{ll} \partial_t u - \frac{1}{\varepsilon} \nabla_{\parallel} \cdot (A_{\parallel} u^{5/2} \nabla_{\parallel} u) - \nabla_{\perp} \cdot (A_{\perp} \nabla_{\perp} u) = 0, & \text{in } [0, T] \times \Omega, \\ \frac{1}{\varepsilon} n_{\parallel} \cdot (A_{\parallel} u^{5/2}(t, \cdot) \nabla_{\parallel} u(t, \cdot)) + n_{\perp} \cdot (A_{\perp} \nabla_{\perp} u(t, \cdot)) = -\gamma u(t, \cdot), & \text{on } [0, T] \times \Gamma_{\perp}, \\ \nabla_{\perp} u(t, \cdot) = 0, & \text{on } [0, T] \times \Gamma_{\parallel}, \\ u(0, \cdot) = u^0(\cdot), & \text{in } \Omega. \end{array} \right. \quad (2.3)$$

The coefficient γ is zero for electrons and $\gamma > 0$ for ions [21, 24]. The problem (2.3) describes the diffusion of an initial temperature u^0 within the time interval $[0, T]$ and its outflow through the boundary Γ_{\perp} . Let us denote in the following the time-space cylinder by $Q_T := (0, T) \times \Omega$. The parameter $0 < \varepsilon \ll 1$ can be very small and is responsible for the high anisotropy of the problem. We shall suppose all along this paper, that the coefficients A_{\parallel} and A_{\perp} are of the same order of magnitude, satisfying

Hypothesis 1. Let Γ_{\perp} consist of two connected components $\Gamma_{in} = \{x \in \Gamma / n \cdot b < 0\}$ and $\Gamma_{out} = \{x \in \Gamma / n \cdot b > 0\}$ such that

either

case A: $n = -b$ on Γ_{in} and $n = b$ on Γ_{out}

or

case B: $n \cdot b > -\varkappa$ on Γ_{in} and $n \cdot b < \varkappa$ on Γ_{out} with some constant $0 < \varkappa < 1$.

All the components Γ_{in} , Γ_{out} and Γ_{\parallel} are sufficiently smooth. We suppose moreover $0 < \varepsilon \leq 1$ and $\gamma \geq 0$ fixed. The diffusion coefficients $A_{\parallel} \in W^{1,\infty}(\bar{\Omega})$ and $A_{\perp} \in \mathbb{M}_{d \times d}(W^{1,\infty}(\bar{\Omega}))$ are

supposed to satisfy

$$0 < A_0 \leq A_{\parallel}(x) \leq A_1, \quad f.a.a. \quad x \in \Omega, \quad (2.4)$$

$$A_0 \|v\|^2 \leq v^t A_{\perp}(x)v \leq A_1 \|v\|^2, \quad \forall v \in \mathbb{R}^d \text{ and } f.a.a. \quad x \in \Omega, \quad (2.5)$$

with $0 < A_0 < A_1$ some constants.

Putting formally $\varepsilon = 0$ in (2.3) leads to the following ill-posed problem, admitting infinitely many solutions

$$\begin{cases} -\nabla_{\parallel} \cdot (A_{\parallel} u^{5/2} \nabla_{\parallel} u) = 0, & \text{in } [0, T] \times \Omega, \\ n_{\parallel} \cdot (A_{\parallel} u^{5/2}(t, \cdot) \nabla_{\parallel} u(t, \cdot)) = 0, & \text{on } [0, T] \times \Gamma_{\perp}, \\ \nabla_{\perp} u(t, \cdot) = 0, & \text{on } [0, T] \times \Gamma_{\parallel}, \\ u(0, \cdot) = u^0(\cdot), & \text{in } \Omega. \end{cases} \quad (2.6)$$

Indeed, all functions which are constant along the field lines, meaning $\nabla_{\parallel} u \equiv 0$, and satisfying moreover the boundary condition on Γ_{\parallel} , are solutions of this problem. From a numerical point of view, this ill-posedness in the limit $\varepsilon \rightarrow 0$ can be detected by the fact, that trying to solve (2.3) with standard schemes leads to a linear system, which is very ill-conditioned for $0 < \varepsilon \ll 1$, in particular with a condition number of the order of $1/\varepsilon$.

The aim of this paper will be to introduce an efficient numerical method, permitting to solve (2.3) accurately on a coarse Cartesian grid, which has not to be adapted to the field lines of b and whose mesh size is independent of the value of ε . The here proposed scheme belongs to the category of Asymptotic-Preserving schemes, meaning they are stable independently of the small parameter ε and consistent with the limit problem, if ε tends to zero. The construction of the here developed AP-scheme is an adaptation of a method introduced by the authors in an elliptic framework (see [7]), to the here considered non-linear and time-dependent problem, and is based on a reformulation of the singularly perturbed problem (2.3) into an equivalent problem, which appears to be well-posed in the limit $\varepsilon \rightarrow 0$. But before introducing the AP-approach, we will start by studying in the following section the mathematical properties of problem (2.3). The test configuration chosen all along this paper is the diffusion of an initial temperature hot spot (see Figure 1) along arbitrary magnetic field lines b .

2.1. Mathematical properties. Before starting with the presentation of the AP numerical scheme, let us first check some properties of the diffusion problem (2.3) for fixed $\varepsilon > 0$.

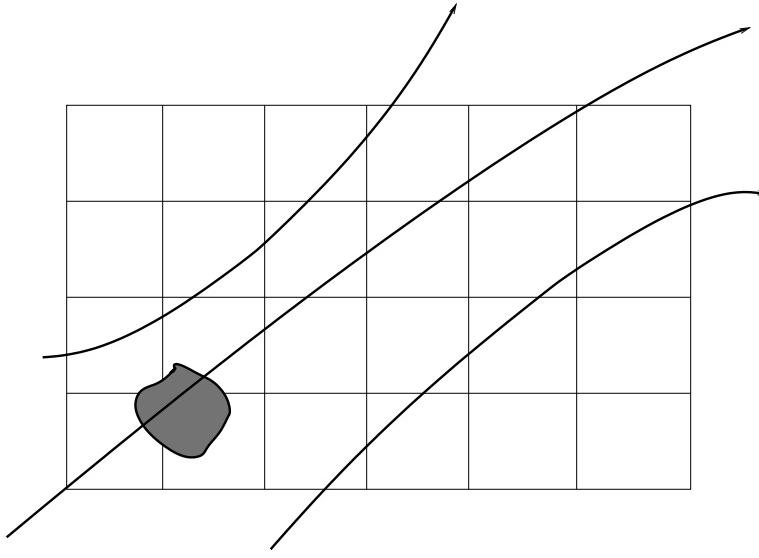


FIGURE 1. Diffusion of a hot temperature spot along the magnetic field lines.

For notational simplicity, we consider a slightly more general form of problem (P)

$$(P_m) \quad \begin{cases} \partial_t u - \nabla_{||} \cdot (A_{||}|u|^{m-1} \nabla_{||} u) - \nabla_{\perp} \cdot (A_{\perp} \nabla_{\perp} u) = 0, & \text{in } [0, T] \times \Omega, \\ A_{||}|u|^{m-1} n_{||} \cdot \nabla_{||} u + A_{\perp} n_{\perp} \cdot \nabla_{\perp} u = -\gamma u, & \text{on } [0, T] \times \Gamma_{\perp}, \\ \nabla_{\perp} u = 0, & \text{on } [0, T] \times \Gamma_{||}, \\ u(0, \cdot) = u^0(\cdot), & \text{in } \Omega, \end{cases} \quad (2.7)$$

for any $m \geq 1$. We obtain the particular case (2.3) by setting $m = 5/2+1$ and redefining $A_{||}$ as $\frac{1}{\varepsilon} A_{||}$ for any $\varepsilon > 0$. Equations of the type (2.7) are rather well studied in the literature. We refer to the classical works [8, 9, 13] as well as to the more modern literature on “The porous medium equation” as reviewed in [1, 22]. However, all these references normally treat only an isotropic version of the problem above, i.e. the non-linearity of the type u^{m-1} is present in front of all the derivatives of u . An anisotropic equation of the form (2.7) is studied in [11], but only in the case $m < \frac{d+1}{d-1}$, so that the value of m pertinent to our application is not covered. Another feature of our setting, which is not sufficiently covered in the existing literature, is the prescription of Robin boundary conditions. This is the reason why we wish to study in this paper the existence, uniqueness and positivity of solutions to (2.7).

We shall first introduce the concept of weak solution of problem (2.7) and state the existence/uniqueness theorem. Note that unlike the literature cited above, we assume from the beginning that the initial conditions are bounded and strictly positive, and prove

the same properties for the weak solutions. Our treatment is thus performed under much less general assumptions than usually required, but this is quite enough for our application.

Definition 2.1. (*Weak solution*) Let $u^0 \in L^\infty(\Omega)$. Then $u \in \mathcal{W}$ with

$$\begin{aligned} \mathcal{W} := \{u \in L^\infty(Q_\infty), \text{ such that } & \forall T > 0 \\ \nabla_\perp u \in L^2(Q_T), \quad |u|^{m-1} \nabla_{||} u \in L^2(Q_T), \quad \partial_t u \in L^2(0, T; (H^1(\Omega))^*)\}, \end{aligned}$$

is called a weak solution of problem (2.7), if $u(0, \cdot) = u^0$ and if for all $T > 0$ one has

$$\begin{aligned} \int_0^T \langle \partial_t u(t, \cdot), \phi(t, \cdot) \rangle_{(H^1)^*, H^1} dt + \int_0^T \int_\Omega A_{||} |u|^{m-1} \nabla_{||} u \cdot \nabla_{||} \phi dxdt \\ + \int_0^T \int_\Omega A_\perp \nabla_\perp u \cdot \nabla_\perp \phi dxdt + \gamma \int_0^T \int_{\Gamma_\perp} u \phi d\sigma dt = 0, \quad \forall \phi \in \mathcal{D} \end{aligned} \quad (2.8)$$

where $\mathcal{D} = L^2(0, T; H^1(\Omega))$.

Remark 2.2. All the terms in this variational formulation are well-defined for any $u \in \mathcal{W}$ and $\phi \in \mathcal{D}$. Indeed, as $u \in L^\infty(Q_T)$, $\nabla_\perp u \in L^2(Q_T)$ and $|u|^{m-1} \nabla_{||} u \in L^2(Q_T)$, one also has $|u|^{m-1} \nabla u \in L^2(Q_T)$, and thus $|u|^m \in L^2(0, T; H^1(\Omega))$. This means that $|u|^m$ has a trace on $\partial\Omega$, belonging to $L^2(\partial\Omega)$. As $L^{2m}(\partial\Omega) \subset L^2(\partial\Omega)$ for all $m \geq 1$ (Ω is bounded), one has $u|_{\partial\Omega} \in L^2(\partial\Omega)$, justifying thus the boundary integral in (2.8). Moreover, we have the continuous inclusion $\mathcal{W} \subset C([0, T]; (H^1(\Omega))^*)$. This shows that one can impose the initial condition $u(0, \cdot) = u^0$ with $u^0 \in L^\infty(\Omega) \subset (H^1(\Omega))^*$.

Remark 2.3. Actually, we have a sharper characterization of continuity in time for functions in \mathcal{W} . Indeed, let Y be a Banach space and let us denote by $C([0, T]; Y_w)$ the space of weakly continuous functions with values in Y , which means for each $\psi \in Y^*$ the mapping $t \mapsto \langle \psi, u(t) \rangle_{Y^*, Y}$ is continuous. Then, one can prove (see [14] for more details) that $\mathcal{W} \subset L^\infty(0, T; L^2(\Omega)) \cap C([0, T]; (H^1(\Omega))^*) \subset C([0, T]; L_w^2(\Omega))$.

Theorem 2.4. (*Existence/Uniqueness/Positivity*) Let $u^0 \in L^\infty(\Omega)$ satisfy $0 < \beta \leq u^0 \leq M < \infty$ on Ω , for some $\beta > 0$. Assume Hypothesis 1 and $m \geq 1$. Then there exists a unique weak solution $u \in \mathcal{W}$ of (2.7), which satisfies $ce^{-Kt} \leq u \leq M$ a.e. on Q_∞ , with some sufficiently small $c > 0$ and some sufficiently large $K > 0$.

Before proving this theorem, let us also define the sub- and super-solutions to problem (2.7) and establish a comparison principle for them.

Definition 2.5. (*Sub/super-solutions*) A function $u \in \mathcal{W}$ is called a weak sub- (resp. super-) solution to problem (2.7) if the variational formulation (2.8) is verified for all $\phi \in \mathcal{D}$ with $\phi \geq 0$ on Q_∞ , and where the equality sign is replaced by \leq (resp. \geq).

Lemma 1. (Comparison principle) Assume Hypothesis 1 and $m \geq 1$. Let u_1 be a non-negative sub-solution and u_2 be a non-negative super-solution to (2.7) such that $u_1(0, x) \leq u_2(0, x)$ for a.a. $x \in \Omega$. If at least one of the functions u_1, u_2 is strictly positive, i.e. $\forall T > 0 \exists \beta_T > 0$ such that $u_1 \geq \beta_T$ or $u_2 \geq \beta_T$ on Q_T , then $u_1 \leq u_2$ on Q_∞ .

Proof. For any $k > 0$, introduce the function $H_k : \mathbb{R} \rightarrow \mathbb{R}$ as

$$H_k(u) = \begin{cases} 0, & \text{if } u \leq 0 \\ ku, & \text{if } 0 < u \leq \frac{1}{k} \\ 1, & \text{if } u > \frac{1}{k} \end{cases}$$

and put $\phi = H_k(u_1^m - u_2^m)$. Note that $\phi \in L^2(0, T; H^1(\Omega))$ since $u_1, u_2 \in \mathcal{W}$ and thus $\nabla u_1^m, \nabla u_2^m \in L^2(Q_T)$. Observe also that the gradient of ϕ is zero outside from the set $\omega_T^k = \{(t, x) \in \bar{Q}_T : 0 < u_1^m - u_2^m < \frac{1}{k}\}$. Choosing this ϕ as the test function in the inequalities (2.8) for u_1 and u_2 and subtracting the second one from the first one gives

$$\begin{aligned} \int_0^T \langle \partial_t(u_1 - u_2), H_k(u_1^m - u_2^m) \rangle_{(H^1)^*, H^1} dt &\leq -\frac{k}{m} \iint_{\omega_T^k} A_{||} |\nabla_{||}(u_1^m - u_2^m)|^2 dxdt \quad (2.9) \\ &- mk \iint_{\omega_T^k} A_{\perp} \nabla_{\perp}(u_1 - u_2) \cdot [u_1^{m-1} \nabla_{\perp}(u_1 - u_2) + (u_1^{m-1} - u_2^{m-1}) \nabla_{\perp} u_2] dxdt \\ &- \gamma \int_0^T \int_{\Gamma_{\perp}} (u_1 - u_2) H_k(u_1^m - u_2^m) d\sigma dt \\ &\leq mk \iint_{\omega_T^k} A_{\perp} |\nabla_{\perp}(u_1 - u_2) \cdot \nabla_{\perp} u_2| (u_1^{m-1} - u_2^{m-1}) dxdt \end{aligned}$$

since $(u_1 - u_2)$ and $H_k(u_1^m - u_2^m)$ are of the same sign. We have moreover

$$u_1^{m-1} - u_2^{m-1} \leq C_m \frac{u_1^m - u_2^m}{u_1 + u_2} \leq \frac{C_m}{k\beta}$$

on ω_T^k with a constant $C_m > 0$ depending only on m . Indeed, the first inequality here holds for any $u_1 \geq u_2 > 0$ and the second inequality follows by noting that $u_1^m - u_2^m < 1/k$ on ω_T^k and $u_1 + u_2 \geq \beta_T$ on Q_T . We see that (2.9) combined with the inequality above implies

$$\int_0^T \langle \partial_t(u_1 - u_2), H_k(u_1^m - u_2^m) \rangle_{(H^1)^*, H^1} dt \leq \frac{mC_m}{\beta_T} \iint_{\omega_T^k} A_{\perp} |\nabla_{\perp}(u_1 - u_2) \cdot \nabla_{\perp} u_2| dxdt.$$

Let us now take the limit $k \rightarrow \infty$ in this inequality. We have $\text{meas}(\omega_T^k) \rightarrow 0$ so that

$$\limsup_{k \rightarrow \infty} \int_0^T \langle \partial_t(u_1 - u_2), H_k(u_1^m - u_2^m) \rangle_{(H^1)^*, H^1} dt \leq 0. \quad (2.10)$$

On the other hand,

$$\lim_{k \rightarrow \infty} \int_0^T \langle \partial_t(u_1 - u_2), H_k(u_1^m - u_2^m) \rangle_{(H^1)^*, H^1} dt = \int_{\Omega} (u_1 - u_2)^+(T, x) dx - \int_{\Omega} (u_1 - u_2)^+(0, x) dx. \quad (2.11)$$

where $(u)^+ = (u + |u|)/2$ denotes the positive part of u . Indeed, $H_k(u_1^m - u_2^m) \rightarrow H(u_1 - u_2)$ a.a. on Q_T where H denotes the Heaviside function ($H(x) = 1$ for $x > 0$ and $H(x) = 0$ for $x \leq 0$). Observing that $\partial_t(u_1 - u_2)^+ = \partial_t(u_1 - u_2)H(u_1 - u_2)$ in the sense of distributions, proves (2.11) for sufficiently smooth u_1, u_2 . A standard density argument shows then that (2.11) actually holds for any $u_1, u_2 \in \mathcal{W}$. Note, in particular, that the terms at the right-hand side of (2.11) are well defined for functions in \mathcal{W} thanks to the inclusion $\mathcal{W} \subset C([0, T]; L_w^2(\Omega))$, cf. Remark 2.3. Comparing (2.10) and (2.11) and taking into account $(u_1 - u_2)^+ = 0$ on Ω at $t = 0$, yields

$$\int_{\Omega} (u_1 - u_2)^+(T, x) dx \leq 0, \quad (2.12)$$

which implies $u_1 \leq u_2$ on Q_{∞} . \square

The construction of the following remarkable sub-solution is essentially due to M. Pierre [19].

Lemma 2. (Construction of a weak solution) *Assume Hypothesis 1 and $m \geq 1$. For any $\beta > 0$, there exists a weak sub-solution w to problem (2.7) satisfying $c \leq w(0, x) \leq \beta$ for $x \in \Omega$ and $w(t, x) \geq ce^{-Kt}$ for $(t, x) \in Q_{\infty}$, with some constants $c, K > 0$ which depend only on β .*

Proof. We will construct a smooth sub-solution w satisfying all the announced properties. We thus rewrite the definition of a sub-solution in the strong form supposing from the beginning that $w \geq 0$:

$$\partial_t w - \frac{1}{m} \nabla_{||} \cdot (A_{||} \nabla_{||} w^m) - \nabla_{\perp} \cdot (A_{\perp} \nabla_{\perp} w) \leq 0, \quad \text{on } (0, \infty) \times \Omega \quad (2.13)$$

$$\frac{1}{m} A_{||} n_{||} \cdot \nabla_{||} w^m + A_{\perp} n_{\perp} \cdot \nabla_{\perp} w + \gamma w \leq 0, \quad \text{on } (0, \infty) \times \Gamma_{\perp}, \quad (2.14)$$

$$n_{\perp} \cdot \nabla_{\perp} w \leq 0, \quad \text{on } (0, \infty) \times \Gamma_{||} \quad (2.15)$$

The construction of such a function w will be performed separately for the two cases mentioned in Hypothesis 1.

Case A: One can construct in this case a new coordinate system ξ_1, \dots, ξ_d on Ω such that the coordinate lines ξ_d coincide with the b -field lines and the surfaces $\xi_d = \text{const}$ are perpendicular to these lines. Domain Ω is represented in these coordinates by a cylinder $\Omega_{\xi} = D \times (0, 1)$ with $\xi' = (\xi_1, \dots, \xi_{d-1}) \in D$ and $\xi_d \in (0, 1)$. We thus have $\nabla_{||} = b\chi \frac{\partial}{\partial \xi_d}$ with some scalar strictly positive field χ . We assume that the component Γ_{in} of the boundary

is represented by $D \times \{\xi_d = 0\}$, Γ_{out} is represented by $D \times \{\xi_d = 1\}$ and $\Gamma_{||}$ is represented by $\partial D \times (0, 1)$.

We are searching now for a sub-solution under the form $w(t, x) = \delta(t)(\sin(\pi\xi_d) + \eta(t))^{1/m}$ where $\delta(t)$ and $\eta(t)$ are two positive decreasing functions which are yet to be adjusted. We observe immediately that $\nabla_{\perp} w = 0$ on Ω for all time so that (2.15) is automatically satisfied. The remaining boundary conditions (2.14) should be checked on Γ_{in} and Γ_{out} . We remind that $n = n_{||} = b$ on Γ_{out} ($\xi_d = 1$). Similarly, $n = n_{||} = -b$ on Γ_{in} ($\xi_d = 0$). Substituting the Ansatz for w into (2.14) now gives

$$-\frac{1}{m} A_{||} \chi \delta^m \pi + \gamma a \eta^{\frac{1}{m}} \leq 0, \text{ for } \xi_d = 0 \text{ and } \xi_d = 1.$$

This holds if one takes $\eta = K_1 \delta^{m(m-1)}$ where $K_1 = \left(\min_{\xi \in \bar{\Omega}_{\xi}} \frac{\pi}{m\gamma} A_{||} \chi \right)^m > 0$.

It remains to check (2.13). Substituting the Ansatz for w , this inequality is reduced to

$$\dot{\delta} (\sin(\pi\xi_d) + \eta)^{\frac{1}{m}} - \frac{\delta^m}{m} \chi \frac{\partial}{\partial \xi_d} (A_{||} \chi) \pi \cos(\pi\xi_d) + \frac{\delta^m}{m} A_{||} \chi^2 \pi^2 \sin(\pi\xi_d) \leq 0. \quad (2.16)$$

Note that we have denoted the time derivative here by a dot and we neglected a term with $\dot{\eta}$ since it is negative (the function $\eta(t)$ is decreasing). We divide now both sides by $(\sin(\pi\xi_d) + \eta)^{\frac{1}{m}}$ and bound each term on the left-hand side as

$$-\frac{\delta^m \pi \cos(\pi\xi_d) \chi}{m (\sin(\pi\xi_d) + \eta)^{\frac{1}{m}}} \frac{\partial(A_{||}\chi)}{\partial \xi_d} \leq \frac{\pi \delta^m \chi}{m \eta^{\frac{1}{m}}} \left| \frac{\partial(A_{||}\chi)}{\partial \xi_d} \right| = \frac{\pi \delta^m \chi}{m K_1^{\frac{1}{m}} \delta^{m-1}} \left| \frac{\partial(A_{||}\chi)}{\partial \xi_d} \right| = \delta \frac{\pi}{m K_1^{\frac{1}{m}}} \chi \left| \frac{\partial(A_{||}\chi)}{\partial \xi_d} \right|$$

and

$$\frac{\delta^m}{m} A_{||} \chi^2 \pi^2 \frac{\sin(\pi\xi_d)}{(\sin(\pi\xi_d) + \eta)^{\frac{1}{m}}} \leq \frac{\delta^m}{m} A_{||} \chi^2 \pi^2 (\sin(\pi\xi_d))^{1-\frac{1}{m}} \leq \frac{\delta^m}{m} A_{||} \chi^2 \pi^2.$$

We see now that inequality (2.16) will be satisfied if we require

$$\dot{\delta} + K_2 \delta + K_3 \delta^m \leq 0 \quad (2.17)$$

with

$$K_2 = \frac{\pi}{m K_1^{\frac{1}{m}}} \max_{\xi \in \Omega_{\xi}} \left| \chi \frac{\partial}{\partial \xi_d} (A_{||} \chi) \right| \text{ and } K_3 = \frac{\pi^2}{m} \max_{\xi \in \Omega_{\xi}} |A_{||} \chi^2|.$$

One can thus take $\delta(t) = \delta_0 e^{-(K_2+K_3)t}$ with any $\delta_0 \in (0, 1]$.

In summary, $w(t, x) = \delta_0 e^{-(K_2+K_3)t} (\sin(\pi\xi_d) + K_1 a_0^{m(m-1)} e^{-m(m-1)(K_2+K_3)t})^{\frac{1}{m}}$ is a sub-solution. Clearly, for any $\beta > 0$ one can take δ_0 small enough so that $w(0, x) \leq \beta$. Moreover, for any t , $w(t, x) \geq \delta_0^m K_1^{1/m} e^{-m(K_2+K_3)t}$ so that we have proved the statement of the Lemma putting $c = \delta_0^m K_1^{1/m}$, $K = m(K_2 + K_3)$.

Case B: Let $\phi \in C^2(\bar{\Omega})$ be a strictly positive function such that $\phi(x) \geq 1$ on $\bar{\Omega}$, $n_{\perp} \cdot \nabla_{\perp} \phi = n \cdot \nabla \phi \leq 0$ on $\Gamma_{||}$ and $\nabla \phi = -\zeta \phi n$ on Γ_{\perp} with some sufficiently big constant $\zeta > 0$, to be prescribed later.

We are searching now for a sub-solution under the form $w(t, x) = ce^{-Kt}\phi(x)$ where c, K are some positive constants which are yet to be adjusted. We observe immediately that (2.15) is automatically satisfied for such w . The left-hand side of (2.14) can be written as

$$\begin{aligned} & \frac{c^m}{m} e^{-mKt} A_{||} n_{||} \cdot \nabla_{||} \phi^m + ce^{-Kt} (A_{\perp} n_{\perp} \cdot \nabla_{\perp} \phi + \gamma \phi) \\ & \leq ce^{-Kt} (-\zeta c^{m-1} e^{-(m-1)Kt} A_{||} |n_{||}|^2 \phi^m - \zeta A_{\perp} n_{\perp} \cdot n_{\perp} \phi + \gamma \phi) \\ & \leq ce^{-Kt} (-\zeta A_{\perp} n_{\perp} \cdot n_{\perp} + \gamma \phi) \end{aligned}$$

and thus it is negative provided ζ is chosen sufficiently big. Indeed, $A_{\perp} n_{\perp} \cdot n_{\perp}$ is uniformly bounded from below by a positive constant in view of the geometrical hypothesis of case B.

It remains to check (2.13). Substituting the Ansatz for w into this inequality yields

$$ce^{-Kt} \left(-K\phi - \frac{c^{m-1} e^{-(m-1)Kt}}{m} \nabla_{||} \cdot (A_{||} \nabla_{||} \phi^m) - \nabla_{\perp} \cdot (A_{\perp} \nabla_{\perp} \phi) \right) \leq 0. \quad (2.18)$$

This inequality is satisfied provided we take $c \leq 1$ and

$$K = \frac{1}{m} \max_{x \in \Omega} |\nabla_{||} \cdot (A_{||} \nabla_{||} \phi^m)| + \max_{x \in \Omega} |\nabla_{\perp} \cdot (A_{\perp} \nabla_{\perp} \phi)|.$$

Finally, for any $\beta > 0$ one can take c small enough so that $w(0, x) = c\phi(x) \leq \beta$. Lemma is thus proved also in case B. \square

Remark 2.6. In the case of a simple “aligned” geometry, i.e. $b = e_d$ and $\Omega = D \times]0, L[$ with D a domain in \mathbb{R}^{d-1} , and supposing $A_{||} = \text{const}$, one can easily construct a sub-solution satisfying a sharper estimate: under the assumptions of the preceding Lemma, there is a sub-solution such that

$$w(t, x) \geq \frac{C}{(1+Kt)^{\frac{m}{m-1}}}.$$

Indeed, one can repeat the proof as in case A of the preceding Lemma, taking $\xi' = (x_1, \dots, x_{d-1})$, $\xi_d = x_d/L$, up to the differential inequality (2.17). One observes now that $K_2 = 0$ so that one can take $\delta(t) = \frac{\delta_0}{(1+Kt)^{\frac{1}{m-1}}}$ with any $\delta_0 > 0$ and $K = (m-1)K_3\delta_0^{m-1}$. Our sub-solution is thus $w = \frac{\delta_0}{(1+Kt)^{\frac{1}{m-1}}} \left(\sin(\pi x_d) + \frac{K_1 \delta_0^{m(m-1)}}{(1+Kt)^m} \right)^{\frac{1}{m}}$ and $w \geq \frac{K_1^{\frac{1}{m}} \delta_0^m}{(1+Kt)^{\frac{m}{m-1}}}$ as stated.

Let us now turn to the proof of our main result.

Proof of Theorem 2.4. We shall first regularize the problem, in order to avoid the degeneracy. Then, in a second step, we shall treat the non-linearity via a fixed point argument. Finally, a priori estimates shall help us to pass to the limit in the regularized

problem, to prove existence. The comparison principle above will be used to establish the uniqueness and the positivity of the solution. Let us now detail these steps.

1st step: Regularization

Fix $0 < \alpha < 1$ and assume that $M > 0$ is an upper bound for u^0 . Introduce for notational simplicity the following functions $a_\alpha, \Lambda_\alpha : \mathbb{R} \rightarrow \mathbb{R}^+$

$$a_\alpha(u) := [\alpha + \min(|u|, M)]^{m-1}, \quad \Lambda_\alpha(u) := \int_0^u a_\alpha(s) ds,$$

and consider the regularized version of (2.8): find $u_\alpha \in W_2^1(0, T; H^1(\Omega), L^2(\Omega))$ (i.e. $u_\alpha \in L^2(0, T; H^1(\Omega))$ and $\partial_t u_\alpha \in L^2(0, T; (H^1(\Omega))^*)$) such that $u_\alpha(0, \cdot) = u^0$ and

$$\begin{aligned} & \int_0^T \langle \partial_t u_\alpha(t, \cdot), \phi(t, \cdot) \rangle_{(H^1)^*, H^1} dt + \int_0^T \int_\Omega A_{||} a_\alpha(u_\alpha) \nabla_{||} u_\alpha \cdot \nabla_{||} \phi dx dt \\ & + \int_0^T \int_\Omega A_\perp \nabla_\perp u_\alpha \cdot \nabla_\perp \phi dx dt + \gamma \int_0^T \int_{\Gamma_\perp} u_\alpha \phi d\sigma dt = 0, \quad \forall \phi \in \mathcal{D}. \end{aligned} \quad (2.19)$$

By standard arguments, this problem is well posed. Indeed, consider the mapping

$$\mathcal{T} : B_R(0) \rightarrow B_R(0), \quad B_R(0) := \{v \in L^2(Q_T) / \|v\|_{L^2(Q_T)} \leq R\},$$

where we associate to any $v \in B_R(0)$ the unique solution $u \in W_2^1(0, T; H^1(\Omega), L^2(\Omega))$ of the linearized, regular parabolic problem

$$\begin{aligned} & \int_0^T \langle \partial_t u(t, \cdot), \phi(t, \cdot) \rangle_{(H^1)^*, H^1} dt + \int_0^T \int_\Omega A_{||} a_\alpha(v) \nabla_{||} u \cdot \nabla_{||} \phi dx dt \\ & + \int_0^T \int_\Omega A_\perp \nabla_\perp u \cdot \nabla_\perp \phi dx dt + \gamma \int_0^T \int_{\Gamma_\perp} u \phi d\sigma dt = 0, \quad \forall \phi \in \mathcal{D}. \end{aligned}$$

Indeed, taking $R := \sqrt{T} \|u^0\|_2$, the mapping \mathcal{T} is well-defined, continuous and $\mathcal{T}(B_R(0))$ is relatively compact in $L^2(Q_T)$. The continuity follows from the fact that for $v_n \rightarrow v$ in $L^2(Q_T)$ and $v_n \rightharpoonup w$ in $W_2^1(0, T; H^1(\Omega), L^2(\Omega))$, the Lebesgue dominated convergence theorem permits us to pass to the limit in the linearized term of the variational formulation. By Schauder fixed point theorem, \mathcal{T} has a fixed point $\mathcal{T}(u) = u$, which provides a solution to (2.19).

The solution u_α of problem (2.19) satisfies $0 \leq u_\alpha \leq M$, provided we have $0 \leq u^0 \leq M$. Indeed, define $u_\alpha^- := \min(0, u_\alpha) \leq 0$. Then one gets for the initial condition $u_\alpha^-(0, \cdot) \equiv 0$. Taking u_α^- as the test function in the variational formulation (2.19), yields immediately

$$\begin{aligned} & \frac{1}{2} \int_\Omega |u_\alpha^-(T, x)|^2 dx + \int_0^T \int_\Omega A_{||} a_\alpha(u_\alpha^-) |\nabla_{||} u_\alpha^-|^2 dx dt \\ & + \int_0^T \int_\Omega A_\perp |\nabla_\perp u_\alpha^-|^2 dx dt + \gamma \int_0^T \int_{\Gamma_\perp} |u_\alpha^-|^2 d\sigma d\tau = 0, \end{aligned}$$

which shows that $u_\alpha^-(T, \cdot) \equiv 0$. Since the same argument can be applied to any final time T , we have $u_\alpha \geq 0$ in Q_∞ .

To prove the estimate from above, define $u_\alpha^+ := \max(0, u_\alpha - M)$. Observe that $u_\alpha^+(0, \cdot) \equiv 0$ and take u_α^+ as the test function in the variational formulation (2.19):

$$\begin{aligned} \frac{1}{2} \int_{\Omega} |u_\alpha^+(T, x)|^2 dx &+ \int_0^T \int_{\Omega} A_{||} a_\alpha(u_\alpha) |\nabla_{||} u_\alpha^+|^2 dx dt \\ &+ \int_0^T \int_{\Omega} A_{\perp} |\nabla_{\perp} u_\alpha^+|^2 dx dt + \gamma \int_0^T \int_{\Gamma_{\perp}} u_\alpha u_\alpha^+ d\sigma d\tau = 0, \end{aligned}$$

which shows that $u_\alpha^+(T, \cdot) \equiv 0$. Since again the same argument can be applied to any final time, we have $u_\alpha \leq M$ in Q_∞ .

2nd step: A priori estimates

In order to pass to the limit $\alpha \rightarrow 0$, we will need some a priori estimates for the solution u_α , independent of α . Taking $\phi = u_\alpha$ in the variational formulation (2.19) yields

$$\begin{aligned} \frac{1}{2} \int_{\Omega} |u_\alpha(T, x)|^2 dx + \int_0^T \int_{\Omega} A_{||} a_\alpha(u_\alpha) |\nabla_{||} u_\alpha|^2 dx dt \\ + \int_0^T \int_{\Omega} A_{\perp} |\nabla_{\perp} u_\alpha|^2 dx dt + \gamma \int_0^T \int_{\Gamma_{\perp}} |u_\alpha|^2 d\sigma dt = \frac{1}{2} \int_{\Omega} |u^0(x)|^2 dx, \end{aligned}$$

which implies

$$\begin{aligned} \|u_\alpha\|_{L^\infty(0,T;L^2(\Omega))} &\leq \|u^0\|_{L^2(\Omega)}, & \int_0^T \int_{\Omega} a_\alpha(u_\alpha) |\nabla_{||} u_\alpha|^2 dx dt &\leq C \|u^0\|_{L^2(\Omega)}^2, \\ \|\nabla_{\perp} u_\alpha\|_{L^2(Q_T)} &\leq C \|u^0\|_{L^2(\Omega)}, & \|u_\alpha\|_{L^2([0,T] \times \Gamma_{\perp})} &\leq C \|u^0\|_{L^2(\Omega)}, \end{aligned} \quad (2.20)$$

with some constant $C > 0$.

Taking now $\phi = \Lambda_\alpha(u_\alpha)$ in (2.19), which is permitted since $u_\alpha \in L^\infty(Q_T) \cap L^2(0, T; H^1(\Omega))$, yields

$$\begin{aligned} \int_0^T \langle \partial_t u_\alpha, \Lambda_\alpha(u_\alpha) \rangle_{(H^1)^*, H^1} dt + \int_0^T \int_{\Omega} A_{||} |\nabla_{||} (\Lambda_\alpha(u_\alpha))|^2 dx dt \\ + \int_0^T \int_{\Omega} A_{\perp} a_\alpha(u_\alpha) |\nabla_{\perp} u_\alpha|^2 dx dt + \gamma \int_0^T \int_{\Gamma_{\perp}} u_\alpha \Lambda_\alpha(u_\alpha) d\sigma d\tau = 0. \end{aligned}$$

The first term can be rewritten as

$$\int_0^T \langle \partial_t u_\alpha, \Lambda_\alpha(u_\alpha) \rangle_{(H^1)^*, H^1} dt = \int_{\Omega} \Psi_\alpha(u_\alpha(T, x)) dx - \int_{\Omega} \Psi_\alpha(u^0(x)) dx,$$

with $\Psi_\alpha(u) := \int_0^u \Lambda_\alpha(s) ds$. Due to the facts that $0 \leq u_\alpha \leq M$, $\Lambda_\alpha(u_\alpha) \geq 0$ and $\Psi_\alpha(u_\alpha) \geq 0$, we get

$$\begin{aligned} \int_0^T \int_{\Omega} A_{||} |\nabla_{||} (\Lambda_\alpha(u_\alpha))|^2 dx dt &\leq C \int_{\Omega} \Psi_\alpha(u^0(x)) dx \leq \frac{C}{m(m+1)} \int_{\Omega} (\alpha + u^0(x))^{m+1} dx \\ &\leq C (\alpha^{m+1} + M^{m+1}). \end{aligned}$$

Thus, we have that the family $\{\nabla_{||}\Lambda_\alpha(u_\alpha)\}_\alpha$, $\alpha \in]0, 1]$ is bounded in $L^2(Q_T)$. Moreover, $\{\nabla_{\perp}\Lambda_\alpha(u_\alpha)\}_\alpha$ is also bounded in $L^2(Q_T)$, since $\nabla_{\perp}(\Lambda_\alpha(u_\alpha)) = a_\alpha(u_\alpha)\nabla_{\perp}u_\alpha$ and $a_\alpha(u_\alpha)$ is uniformly bounded by some positive constant. Hence, $\{\Lambda_\alpha(u_\alpha)\}_\alpha$ is bounded in $L^2(0, T; H^1(\Omega))$.

Let $V = H^1(\Omega) \cap L^\infty(\Omega)$ be the Banach space with the norm $\|\cdot\|_V = \|\cdot\|_{H^1(\Omega)} + \|\cdot\|_{L^\infty(\Omega)}$. For any ϕ in $L^\infty(0, T; V)$,

$$\left| \int_0^T \langle \partial_t \Lambda_\alpha(u_\alpha), \phi \rangle_{(H^1)^*, H^1} dt \right| = \left| \int_0^T \langle \partial_t u_\alpha, a_\alpha(u_\alpha) \phi \rangle_{(H^1)^*, H^1} dt \right| \leq C \|\phi\|_{L^\infty(0, T; V)}$$

with a constant C independent of α . This follows from the variational formulation (2.19) with ϕ replaced by $a_\alpha(u_\alpha)\phi$ and from the estimates (2.20). We see thus that the family $\{\partial_t \Lambda_\alpha(u_\alpha)\}_\alpha$ is bounded in $L^1(0, T; V^*)$. We remind also that $\{\Lambda_\alpha(u_\alpha)\}_\alpha$ is bounded in $L^2(0, T; H^1(\Omega))$. Aubin-Simon compactness lemma [20] applied to the triple of spaces $H^1(\Omega) \subset L^2(\Omega) \subset V^*$ allows us now to conclude that the set $\{\Lambda_\alpha(u_\alpha)\}_\alpha$ is relatively compact in $L^2(0, T; L^2(\Omega)) = L^2(Q_T)$.

3rd step: Passage to the limit

The aim now is to pass to the limit $\alpha \rightarrow 0$ in the variational formulation (2.19) in order to show the existence of a weak solution of problem (2.3). The a priori estimates of the last step permit us to show, that there exists a function $u \in L^2(Q_T)$ satisfying $0 \leq u \leq M$ in Q_T and such that after extracting a sub-sequence from $\{u_\alpha\}_\alpha$, we have

$$\begin{aligned} u_\alpha &\rightharpoonup_{\alpha \rightarrow 0} u \quad \text{in } L^2(Q_T), \quad u_\alpha|_{\Gamma_\perp} &\rightharpoonup_{\alpha \rightarrow 0} u|_{\Gamma_\perp} \quad \text{in } L^2([0, T] \times \Gamma_\perp), \\ \nabla_{\perp} u_\alpha &\rightharpoonup_{\alpha \rightarrow 0} \nabla_{\perp} u \quad \text{in } L^2(Q_T) \quad \text{and} \quad \partial_t u_\alpha &\rightharpoonup_{\alpha \rightarrow 0} \partial_t u \quad \text{in } L^2(0, T; (H^1(\Omega))^*). \end{aligned}$$

To pass to the limit in the non-linear term, we use first the fact that $\{\Lambda_\alpha(u_\alpha)\}_\alpha$ is bounded in $L^2(0, T; H^1(\Omega))$, so that there exists some function $w \in L^2(0, T; H^1(\Omega))$ satisfying

$$\Lambda_\alpha(u_\alpha) \rightharpoonup_{\alpha \rightarrow 0} w \quad \text{in } L^2(0, T; H^1(\Omega)).$$

In order to identify the function w , we need some pointwise convergence of the sequence u_α . For this, recall that the sequence $\{\Lambda_\alpha(u_\alpha)\}_\alpha$ is relatively compact in $L^2(Q_T)$. Thus, up to a sub-sequence

$$\Lambda_\alpha(u_\alpha) \rightarrow_{\alpha \rightarrow 0} w \quad \text{in } L^2(Q_T), \quad \text{hence} \quad \Lambda_\alpha(u_\alpha) \rightarrow_{\alpha \rightarrow 0} w, \quad \text{a.e. in } Q_T.$$

Since for any fixed $u \in [0, M]$, $\Lambda_\alpha(u) \rightarrow_{\alpha \rightarrow 0} \Lambda(u) := \frac{1}{m}u^m$, we have $u_\alpha \rightarrow_{\alpha \rightarrow 0} \Lambda^{-1}w$ a.e. in Q_T . This permits to identify the function $\Lambda^{-1}w$ with u , so that $w = \frac{1}{m}u^m \in L^2(0, T; H^1(\Omega))$.

All these convergences allow now to pass to the limit in the variational formulation (2.19) and to conclude the proof of existence for problem (2.7).

4th step: Positivity and uniqueness

Let u be a weak solution to (2.7). We use first the sub-solution constructed in Lemma 2 and the comparison principle in Lemma 1 to verify that $u \geq ce^{-Kt}$ with some positive

constants c small enough and K large enough. We remark then that M is a super-solution to (2.7). Again, by Lemma 1 we see that $u \leq M$.

Suppose now that (2.8) admits two solutions u_1 and u_2 in \mathcal{W} with the same initial condition $u_1 = u_2 = u^0$ at $t = 0$. We know already that they are both strictly positive, so that Lemma 1 implies $u_1 \leq u_2$ on Q_∞ . Since the two solutions u_1 and u_2 are perfectly interchangeable in the above argument, we have also $u_1 \geq u_2$ and thus $u_1 = u_2$ on Q_∞ . \square

Let us remark here, that due to the strict positivity of the solution, in particular to the property that $u \geq ce^{-Kt}$ a.e. on Q_∞ , with some $c > 0$ and $K > 0$, we have

Corollary 2.7. *Under the hypothesis of theorem 2.4, the weak solution u of (2.7) belongs to the more regular space*

$$\tilde{\mathcal{W}} := \{u \in L^\infty(Q_\infty), \text{ such that } \forall T > 0 : u \in L^2(0, T; H^1(\Omega)), \partial_t u \in L^2(0, T; (H^1(\Omega))^*)\}.$$

Moreover, one has $\frac{d}{dt} \|u(t, \cdot)\|_{L^2(\Omega)} \leq 0$.

3. NUMERICAL METHOD

3.1. Semi-discretization in space. The singular perturbation problem (2.3) is a highly anisotropic equation. Its variational formulation reads: find $u(t, \cdot) \in \mathcal{V} := H^1(\Omega)$ such that

$$(P) \quad \begin{aligned} \langle \partial_t u(t, \cdot), v \rangle_{\mathcal{V}^*, \mathcal{V}} + \frac{1}{\varepsilon} \int_{\Omega} A_{||} |u|^{5/2} \nabla_{||} u(t, \cdot) \cdot \nabla_{||} v \, dx \\ + \int_{\Omega} A_{\perp} \nabla_{\perp} u(t, \cdot) \cdot \nabla_{\perp} v \, dx + \gamma \int_{\Gamma_{\perp}} u(t, \cdot) v \, d\sigma = 0, \quad \forall v \in \mathcal{V} \end{aligned} \quad (3.21)$$

for almost every $t \in (0, T)$. As mentioned already in Section 2, this problem becomes ill-posed if we take formally the limit $\varepsilon \rightarrow 0$. Indeed, only the leading term survives in this limit, so that any function from the space

$$\mathcal{G} := \{p \in \mathcal{V} / \nabla_{||} p = 0 \text{ in } \Omega\}$$

would be a solution. It is easy to establish, however, the well-posed problem for the limit of the solutions to (P) as $\varepsilon \rightarrow 0$. Indeed, one can restrain the test functions in (P) to be in the space \mathcal{G} so that the ε -dependent term disappears and the correct problem in the limit $\varepsilon \rightarrow 0$ reads: find $u(t, \cdot) \in \mathcal{G}$ such that

$$(L) \quad \langle \partial_t u(t, \cdot), v \rangle_{\mathcal{V}^*, \mathcal{V}} + \int_{\Omega} A_{\perp} \nabla_{\perp} u(t, \cdot) \cdot \nabla_{\perp} v \, dx + \gamma \int_{\Gamma_{\perp}} u(t, \cdot) v \, d\sigma = 0, \quad \forall v \in \mathcal{G}$$

for almost every $t \in (0, T)$.

The discussion above shows that a straight-forward discretization of problem (P) may lead to very inaccurate results when $\varepsilon \ll 1$. Indeed, setting $\varepsilon = 0$ would result in a singular problem, so that the problem with $\varepsilon \ll 1$ would be very ill-conditioned. To cope with this difficulty and to obtain a numerical scheme which is uniformly accurate with respect to ε ,

we introduce an Asymptotic-Preserving reformulation, very similar to the one introduced in [7]. The idea is to rewrite the singularly perturbed problem (3.21) in an equivalent form, which is however well-posed when one sets there formally $\varepsilon = 0$ and gives moreover the correct limit problem (L). In order to do this, we introduce the auxiliary unknown q by the relation $\varepsilon \nabla_{\parallel} q = u^{5/2} \nabla_{\parallel} u$ in Ω and $q = 0$ on Γ_{in} , which rescales the nasty part of the equation permitting to get rid of the terms of order $O(1/\varepsilon)$. The reformulated problem, called in the sequel the Asymptotic-Preserving reformulation (AP-problem) reads: find $(u(t, \cdot), q(t, \cdot)) \in \mathcal{V} \times \mathcal{L}$, solution of

$$(AP) \quad \begin{cases} \langle \frac{\partial u}{\partial t}, v \rangle_{\mathcal{V}^*, \mathcal{V}} + \int_{\Omega} (A_{\perp} \nabla_{\perp} u) \cdot \nabla_{\perp} v \, dx + \int_{\Omega} A_{\parallel} \nabla_{\parallel} q \cdot \nabla_{\parallel} v \, dx + \gamma \int_{\Gamma_{\perp}} uv \, ds = 0, \\ \int_{\Omega} A_{\parallel} u^{5/2} \nabla_{\parallel} u \cdot \nabla_{\parallel} w \, dx - \varepsilon \int_{\Omega} A_{\parallel} \nabla_{\parallel} q \cdot \nabla_{\parallel} w \, dx = 0, \quad \forall w \in \mathcal{L}, \end{cases} \quad (3.22)$$

where

$$\mathcal{L} := \{q \in L^2(\Omega) / \nabla_{\parallel} q \in L^2(\Omega) \text{ and } q|_{\Gamma_{in}} = 0\}. \quad (3.23)$$

System (3.22) is an equivalent reformulation (for fixed $\varepsilon > 0$) of the original P-problem (3.21). Putting now formally $\varepsilon = 0$ in (AP) leads to the well-posed limit problem

$$(L') \quad \begin{cases} \langle \frac{\partial u}{\partial t}, v \rangle_{\mathcal{V}^*, \mathcal{V}} + \int_{\Omega} (A_{\perp} \nabla_{\perp} u) \cdot \nabla_{\perp} v \, dx + \int_{\Omega} A_{\parallel} \nabla_{\parallel} q \cdot \nabla_{\parallel} v \, dx + \gamma \int_{\Gamma_{\perp}} uv \, ds = 0, \\ \int_{\Omega} A_{\parallel} u^{5/2} \nabla_{\parallel} u \cdot \nabla_{\parallel} w \, dx = 0, \quad \forall w \in \mathcal{L}, \end{cases} \quad (3.24)$$

which is equivalent to problem (L). Note that q acts here as a Lagrange multiplier for the constraint $u \in \mathcal{G}$, which provides the uniqueness of the solution. Hence the AP-reformulation permits a continuous transition from the *P*-model to the *L*-model, which enables the uniform accuracy of the scheme with respect to ε .

Let us now choose a triangulation of the domain Ω with triangles or quadrangles of order h and introduce the finite element spaces $\mathcal{V}_h \subset \mathcal{V}$ and $\mathcal{L}_h \subset \mathcal{L}$ of type \mathbb{P}_k or \mathbb{Q}_k on this mesh. The finite element discretization of (3.22) writes then: find $(u_h, q_h) \in \mathcal{V}_h \times \mathcal{L}_h$ such that

$$(AP)_h \quad \begin{cases} \int_{\Omega} \frac{\partial u_h}{\partial t} v_h \, dx + \int_{\Omega} (A_{\perp} \nabla_{\perp} u_h) \cdot \nabla_{\perp} v_h \, dx + \int_{\Omega} A_{\parallel} \nabla_{\parallel} q_h \cdot \nabla_{\parallel} v_h \, dx + \gamma \int_{\Gamma_{\perp}} u_h v_h \, ds = 0, \\ \int_{\Omega} A_{\parallel} u_h^{5/2} \nabla_{\parallel} u_h \cdot \nabla_{\parallel} w_h \, dx - \varepsilon \int_{\Omega} A_{\parallel} \nabla_{\parallel} q_h \cdot \nabla_{\parallel} w_h \, dx = 0, \quad \forall w_h \in \mathcal{V}_h. \end{cases} \quad (3.25)$$

Remark that this system is continuous in time and also nonlinear, such that one has to develop now a procedure for the linearization and the discretization in time. This procedure has to be chosen carefully, such that the AP-property developed so far, is not destroyed. This is the aim of the next section.

3.2. Semi-discretization in time. In order to approach numerically the time derivative in (3.25), we introduce three different schemes : a standard first order, implicit Euler scheme, the Crank-Nicolson scheme and a second order, L-stable Runge-Kutta method. We show in the following that the first order Euler-scheme is stable and asymptotic-preserving. The Crank-Nicolson schemes gives reliable results and second order convergence under certain assumptions, but is not asymptotic-preserving. Thus, if second order accuracy in time is desired, the L-stable Runge-Kutta method has to be applied. All three methods are exposed to numerical tests and compared in Section 4.

3.2.1. Implicit Euler scheme.

Introducing the forms

$$(\Theta, \chi) := \int_{\Omega} \Theta \chi \, dx, \quad (3.26)$$

$$a_{\parallel nl}(\Psi, \Theta, \chi) := \int_{\Omega} A_{\parallel} \Psi^{5/2} \nabla_{\parallel} \Theta \cdot \nabla_{\parallel} \chi \, dx, \quad (3.27)$$

$$a_{\parallel}(\Theta, \chi) := \int_{\Omega} A_{\parallel} \nabla_{\parallel} \Theta \cdot \nabla_{\parallel} \chi \, dx, \quad a_{\perp}(\Theta, \chi) := \int_{\Omega} A_{\perp} \nabla_{\perp} \Theta \cdot \nabla_{\perp} \chi \, dx, \quad (3.28)$$

allows us to write the first order, implicit Euler method in the compact notation: Find $(u_h^{n+1}, q_h^{n+1}) \in \mathcal{V}_h \times \mathcal{L}_h$, solution of

$$(E_{AP}) \quad \begin{cases} (u_h^{n+1}, v_h) + \tau \left(a_{\perp}(u_h^{n+1}, v_h) + a_{\parallel}(q_h^{n+1}, v_h) + \gamma \int_{\Gamma_{\perp}} u_h^{n+1} v_h \, ds \right) = (u_h^n, v_h) \\ a_{\parallel nl}(u_h^n, u_h^{n+1}, w_h) - \varepsilon a_{\parallel}(q_h^{n+1}, w_h) = 0, \end{cases}, \quad (3.29)$$

where the non linear term $(u_h^{n+1})^{5/2}$ was replaced by a first order approximation in τ :

$$(u_h^{n+1})^{5/2} = (u_h^n + O(\tau))^{5/2} = (u_h^n)^{5/2} + O(\tau). \quad (3.30)$$

A slightly different first order AP-scheme was introduced in [16] for the resolution of the same temperature balance problem. There, the (P)-problem was firstly discretized in time (implicit Euler), then linearized by a fixed point mapping, and finally the AP reformulation applied. The numerical results obtained in [16] are similar to the present ones.

3.2.2. Linearized Crank-Nicolson scheme.

To construct a scheme, which is second order in time, one can come to the idea to employ the Crank-Nicolson scheme: Find $(u_h^{n+1}, q_h^{n+1}) \in \mathcal{V}_h \times \mathcal{L}_h$, solution of

$$\begin{cases} (u_h^{n+1}, v_h) + \tau \left(a_{\perp}(u_h^{n+1/2}, v_h) + a_{\parallel}(q_h^{n+1}, v_h) + \gamma \int_{\Gamma_{\perp}} u_h^{n+1/2} v_h \, ds \right) = (u_h^n, v_h) \\ a_{\parallel nl}(u_h^{n+1/2}, u_h^{n+1/2}, w_h) - \varepsilon a_{\parallel}(q_h^{n+1}, w_h) = 0. \end{cases} \quad (3.31)$$

As one can observe, we have to deal for each fixed n , with a nonlinear equation. In the linear terms, one can set $u_h^{n+1/2} = \frac{1}{2}(u_h^{n+1} + u_h^n)$. To linearize the term $a_{\parallel nl}(u_h^{n+1/2}, u_h^{n+1/2}, w_h)$ however, we shall use the standard linear extrapolation method. In other words, the nonlinearity in this last formula, $(u_h^{n+1/2})^{5/2}$, will be replaced by a linearized second order approximation in τ :

$$(u_h^{n+1/2})^{5/2} = \left(u_h^n + \frac{1}{2}(u_h^n - u_h^{n-1}) + O(\tau^2) \right)^{5/2} = \left(u_h^n + \frac{1}{2}(u_h^n - u_h^{n-1}) \right)^{5/2} + O(\tau^2). \quad (3.32)$$

The resulting linear system reads finally: Find $(u_h^{n+1}, q_h^{n+1}) \in \mathcal{V}_h \times \mathcal{L}_h$, solution of

$$(CN_{AP}) \begin{cases} (u_h^{n+1}, v_h) + \frac{\tau}{2} \left(a_{\perp}(u_h^{n+1}, v_h) + \gamma \int_{\Gamma_{\perp}} u_h^{n+1} v_h \, ds \right) + \tau a_{\parallel}(q_h^{n+1}, v_h) \\ \quad = (u_h^n, v_h) - \frac{\tau}{2} \left(a_{\perp}(u_h^n, v_h) + \gamma \int_{\Gamma_{\perp}} u_h^n v_h \, ds \right), \\ \frac{1}{2} a_{\parallel nl} \left(\frac{1}{2} (3u_h^n - u_h^{n-1}), u_h^{n+1}, w_h \right) - \varepsilon a_{\parallel}(q_h^{n+1}, w_h) \\ \quad = -\frac{1}{2} a_{\parallel nl} \left(\frac{1}{2} (3u_h^n - u_h^{n-1}), u_h^n, w_h \right). \end{cases} \quad (3.33)$$

Unfortunately, this method is not Asymptotic-Preserving. For small values of ε one expects that the solution will immediately fall into the space of functions almost constant in the direction of the anisotropy, no matter what initial condition was imposed. In the case of the Crank-Nicolson scheme for large time steps compared to $\varepsilon/(u_h^n)^{5/2}$, the second equation in (3.33) will constrain the numerical solution to oscillate if the initial condition is not already in the suitable space. This requires the restrictive choice of a time step of the order of $\varepsilon/(u_h^n)^{5/2}$, which yields the method inapplicable in general cases. In other words, the Crank-Nicolson scheme is unable to model diffusion processes for large Δt , due to the inadequate approximation of the damping processes. The Crank-Nicolson scheme is A-stable but not L-stable and the AP-property of a scheme is strongly related to the L-stability of the scheme.

As an example of the non-convergence of the (CN_{AP}) scheme in a general case, we show some numerical results corresponding to a test case defined in the Section 4.2.2. The initial condition is a Gaussian peak located in the center of the computational domain with a maximum of $10^5 K$. If the time step τ is too large, than u_h^n will immediately reach negative values and thus the numerical algorithm will fail in the next iteration. However,

if τ is sufficiently small the (CN_{AP}) scheme is of second order in time. Unfortunately, the biggest time step that does not provoke oscillations in the numerical solution, is of the order of $10^{-16}s$, for an initial Gaussian peak of 10^5K . This makes the (CN_{AP}) scheme of no practical use in real simulations. These results are plotted on Figure 2.

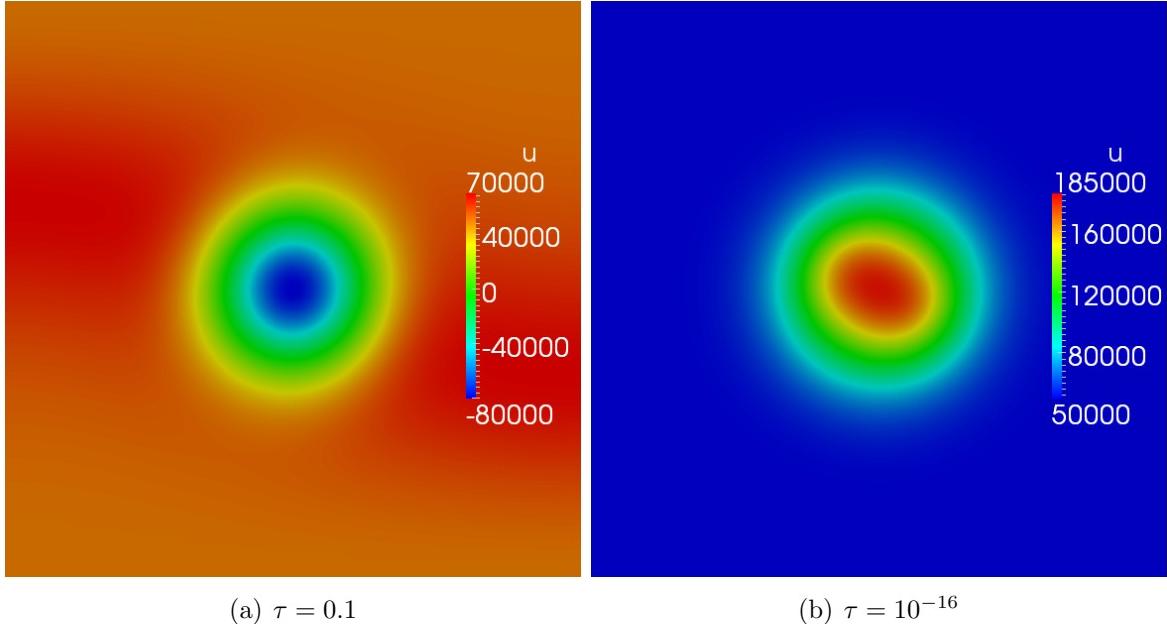


FIGURE 2. Non convergence of the (CN_{AP}) scheme. Negative values of u_h^n are obtained after one iteration of the method, for big time steps. If the time step is sufficiently small, the method converges.

3.2.3. *L-stable Runge-Kutta method.*

As we are interested in an AP-scheme, which is second order accurate in time, we propose now a two stage Diagonally Implicit Runge-Kutta (DIRK) second order scheme, which does not suffer from the limitations of the Crank-Nicolson discretization. The scheme is developed according to the following Butcher's diagram:

$$\begin{array}{c|cc} \lambda & \lambda & 0 \\ \hline 1 & 1-\lambda & \lambda \\ \hline & 1-\lambda & \lambda \end{array} \quad (3.34)$$

with $\lambda = 1 - \frac{1}{\sqrt{2}}$.

Remark 3.1. (*Butcher's diagram*) The coefficients of the s -stage Runge-Kutta method are usually displayed in a Butcher's diagram :

$$\begin{array}{c|ccc} c_1 & a_{11} & \cdots & a_{1s} \\ \vdots & \vdots & & \vdots \\ c_s & a_{s1} & \cdots & a_{ss} \\ \hline & b_1 & \cdots & b_s \end{array} \quad (3.35)$$

Applying this method to approximate to following problem

$$\frac{\partial u}{\partial t} = Lu + f(t), \quad (3.36)$$

reads: For given u^n , being an approximation of $u(t_n)$, the u^{n+1} is determined accordingly to :

$$u_i = u^n + \tau \sum_{j=1}^s a_{ij} (Lu_j + f(t + c_j \tau)), \quad (3.37)$$

$$u^{n+1} = u^n + \sum_{j=1}^s b_j u_j. \quad (3.38)$$

If $b_j = a_{sj}$ for $j = 1, \dots, s$ than $u^{n+1} = u_s$.

The scheme (3.34) is known to be L-stable, thus providing the Asymptotic Preserving property. The scheme writes: Find $(u_h^{n+1}, q_h^{n+1}) \in \mathcal{V}_h \times \mathcal{L}_h$, solution of

$$(RK_{AP}) \quad \left\{ \begin{array}{l} (u_{1,h}^{n+1}, v_h) + \tau \lambda \left(a_{\perp}(u_{1,h}^{n+1}, v_h) + \gamma \int_{\Gamma_{\perp}} u_{1,h}^{n+1} v_h ds + a_{\parallel}(q_{1,h}^{n+1}, v_h) \right) \\ \quad = (u_h^n, v_h) \\ a_{\parallel nl} (u_h^n + \lambda(u_h^n - u_h^{n-1}), u_{1,h}^{n+1}, w_h) - \varepsilon a_{\parallel}(q_{1,h}^{n+1}, w_h) = 0 \\ \\ (u_{2,h}^{n+1}, v_h) + \tau \lambda \left(a_{\perp}(u_{2,h}^{n+1}, v_h) + \gamma \int_{\Gamma_{\perp}} u_{2,h}^{n+1} v_h ds + a_{\parallel}(q_{2,h}^{n+1}, v_h) \right) \\ \quad = (u_h^n, v_h) + \frac{1-\lambda}{\lambda} (u_{1,h}^{n+1} - u_h^n, v_h) \\ a_{\parallel nl} (u_h^n + (u_h^n - u_h^{n-1}), u_{2,h}^{n+1}, w_h) - \varepsilon a_{\parallel}(q_{2,h}^{n+1}, w_h) = 0 \end{array} \right. \quad (3.39)$$

$$u_h^{n+1} = u_{2,h}^{n+1}, \quad q_h^{n+1} = q_{2,h}^{n+1},$$

with $u_{1,h}^{n+1}$ (respectively $u_{2,h}^{n+1}$) being the solution of the first (respectively second) stage of the Runge-Kutta method. The terms $u_h^n + \lambda(u_h^n - u_h^{n-1})$ and $u_h^n + (u_h^n - u_h^{n-1})$ are respectively the second order time-approximations of $u_h(t + \lambda\tau)$ and $u_h(t + \tau)$, used to linearize the problem.

For each time step we have therefore to assemble and solve two linearized problems. This method is two times slower than the Crank-Nicolson scheme, with the advantage however of maintaining the AP-property of the scheme, advantage which is crucial for $0 < \varepsilon \ll 1$.

4. NUMERICAL RESULTS

In this section we compare the proposed implicit Euler-AP and DIRK-AP schemes with a standard linearized implicit Euler discretization of the initial singular perturbation problem (2.3), given by

$$(P)_{h\tau} \quad (u_h^{n+1}, v_h) + \tau \left(a_\perp(u_h^{n+1}, v_h) + \frac{1}{\varepsilon} a_{\parallel nl}(u_h^n, u_h^{n+1}, v_h) + \gamma \int_{\Gamma_\perp} u_h^{n+1} v_h \, ds \right) = (u_h^n, v_h). \quad (4.40)$$

4.1. Discretization. Let us present the space discretization in a 2D case. We consider a square computational domain $\Omega = [0, 1] \times [0, 1]$. All simulations are performed on structured meshes. Let us introduce the Cartesian, homogeneous grid

$$x_i = i/N_x, \quad 0 \leq i \leq N_x, \quad y_j = j/N_y, \quad 0 \leq j \leq N_y, \quad (4.41)$$

where N_x and N_y are positive even constants, corresponding to the number of discretization intervals in the x - resp. y -direction. The corresponding mesh-sizes are denoted by $h_x > 0$ resp. $h_y > 0$. Choosing a \mathbb{Q}_2 finite element method (\mathbb{Q}_2 -FEM), based on the following quadratic base functions

$$\theta_{x_i} = \begin{cases} \frac{(x-x_{i-2})(x-x_{i-1})}{2h_x^2} & x \in [x_{i-2}, x_i], \\ \frac{(x_{i+2}-x)(x_{i+1}-x)}{2h_x^2} & x \in [x_i, x_{i+2}], \\ 0 & \text{else} \end{cases}, \quad , \quad \theta_{y_j} = \begin{cases} \frac{(y-y_{j-2})(y-y_{j-1})}{2h_y^2} & y \in [y_{j-2}, y_j], \\ \frac{(y_{j+2}-y)(y_{j+1}-y)}{2h_y^2} & y \in [y_j, y_{j+2}], \\ 0 & \text{else} \end{cases}, \quad (4.42)$$

for even i, j and

$$\theta_{x_i} = \begin{cases} \frac{(x_{i+1}-x)(x-x_{i-1})}{h_x^2} & x \in [x_{i-1}, x_{i+1}], \\ 0 & \text{else} \end{cases}, \quad , \quad \theta_{y_j} = \begin{cases} \frac{(y_{j+1}-y)(y-y_{j-1})}{h_y^2} & y \in [y_{j-1}, y_{j+1}], \\ 0 & \text{else} \end{cases}, \quad (4.43)$$

for odd i, j , we define the space

$$W_h := \{v_h = \sum_{i,j} v_{ij} \theta_{x_i}(x) \theta_{y_j}(y)\}.$$

The spaces \mathcal{V}_h and \mathcal{L}_h are then defined by

$$\mathcal{V}_h = \mathcal{W}_h, \quad \mathcal{L}_h = \{q_h \in \mathcal{V}_h, \text{ such that } q_h|_{\Gamma_{in}} = 0\}.$$

The matrix elements are computed using the 2D Gauss quadrature formula, with 3 points in the x and y direction:

$$\int_{-1}^1 \int_{-1}^1 f(\xi, \eta) d\xi d\eta = \sum_{i,j=-1}^1 \omega_i \omega_j f(\xi_i, \eta_j), \quad (4.44)$$

where $\xi_0 = \eta_0 = 0$, $\xi_{\pm 1} = \eta_{\pm 1} = \pm \sqrt{\frac{3}{5}}$, $\omega_0 = 8/9$ and $\omega_{\pm 1} = 5/9$, which is exact for polynomials of degree 5. Linear systems obtained for all methods in these numerical experiments are solved using a LU decomposition, implemented by the MUMPS library.

4.2. Numerical tests.

4.2.1. Known analytical solution.

First, let us construct a numerical test case with a known solution. Finding an analytical solution for an arbitrary b -field presents a considerable difficulty. In the previous paper [6], we presented a way to find such a solution. Let us recall briefly how to do this. The starting point is a limit solution

$$u^0 = (\cos(\pi y + \alpha(y^2 - y) \cos(\pi x)) + 4) T_m e^{-t}, \quad (4.45)$$

where α is a numerical constant aimed to control the variations of b . For $\alpha = 0$, the limit solution represents a solution for the constant b case. The parameter T_m is the scaling of u_0 .

Since u^0 is a limit solution, it is constant along the b field lines. Therefore we can determine the b field using the following implication

$$\nabla_{\parallel} u^0 = 0 \Rightarrow b_x \frac{\partial u^0}{\partial x} + b_y \frac{\partial u^0}{\partial y} = 0, \quad (4.46)$$

which yields for example

$$b = \frac{B}{|B|}, \quad B = \begin{pmatrix} \alpha(2y - 1) \cos(\pi x) + \pi \\ \pi \alpha(y^2 - y) \sin(\pi x) \end{pmatrix}. \quad (4.47)$$

Note that the field B , constructed in this way, satisfies $\operatorname{div} B = 0$, which is an important property in the framework of plasma simulations. Furthermore, we have $B \neq 0$ in the computational domain. Now, we choose u^ε to be a function that converges, as $\varepsilon \rightarrow 0$, to the limit solution u^0 , for example

$$p = (\cos(\pi y + \alpha(y^2 - y) \cos(\pi x)) + 4) T_m e^{-t} \quad (4.48)$$

$$q = p^{-3/2} \sin(3\pi x)/3\pi \quad (4.49)$$

$$u = p + \varepsilon q. \quad (4.50)$$

In our simulations we set $\alpha = 1$ so that the direction of the anisotropy is variable in the computational domain. Note that we have

$$-\frac{1}{\varepsilon} u^{5/2}(t, 0) \nabla_{\parallel} u(t, 0) = u(t, 0), \quad (4.51)$$

$$\frac{1}{\varepsilon} u^{5/2}(t, 1) \nabla_{\parallel} u(t, 1) = u(t, 1). \quad (4.52)$$

The problem is supplied with a force term computed accordingly.

As an initial condition we take $u(t = 0)$, with u defined by (4.50). In this setting we expect both Asymptotic-Preserving methods (E_{AP}) and (RK_{AP}) to converge in the optimal rate, independently on ε and b .

First we test the space convergence of the methods. To do this we choose a small time step such that the time discretization error is much smaller than the space discretization error. We then vary the mesh size and perform simulations for 100 time steps. The results are summarized in Table 1 and Figure 3. All three methods give as expected the third order space convergence in the L_2 -norm for large values of ε . Moreover, due to the extremely small time step, the numerical precision is the same, even if one uses second or first order methods. For small values of ε only the Asymptotic Preserving schemes give good numerical solutions.

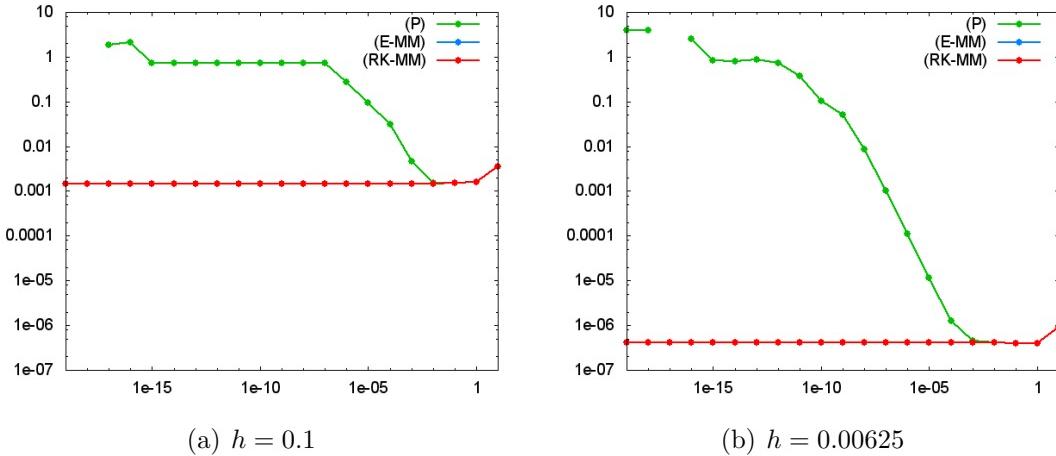


FIGURE 3. Relative L^2 -errors between the exact solution u^ε and the computed solution for the standard scheme (P), Euler AP method (E_{AP}) and DIRK AP scheme (RK_{AP}) as a function of ε and for $h = 0.1$ resp.
 $h = 0.00625$. The time step is $\tau = 10^{-6}$.

Finally we test the time convergence of the methods. To do this we choose a small mesh size such that the space discretization error is smaller than the time discretization error. We then vary the time step and perform simulations on a fixed grid. The results are

h	L^2 -error $\varepsilon = 1$		
	P	E_{AP}	RK_{AP}
0.1	1.60×10^{-3}	1.60×10^{-3}	1.60×10^{-3}
0.05	2.02×10^{-4}	2.02×10^{-4}	2.02×10^{-4}
0.025	2.55×10^{-5}	2.55×10^{-5}	2.55×10^{-5}
0.0125	3.2×10^{-6}	3.2×10^{-6}	3.2×10^{-6}
0.00625	4.0×10^{-7}	4.0×10^{-7}	4.0×10^{-7}

h	L^2 -error $\varepsilon = 10^{-10}$		
	P	E_{AP}	RK_{AP}
0.1	7.3×10^{-1}	1.47×10^{-3}	1.47×10^{-3}
0.05	7.3×10^{-1}	2.04×10^{-4}	2.04×10^{-4}
0.025	7.3×10^{-1}	2.65×10^{-5}	2.65×10^{-5}
0.0125	4.9×10^{-1}	3.3×10^{-6}	3.3×10^{-6}
0.00625	1.04×10^{-1}	4.2×10^{-7}	4.2×10^{-7}

TABLE 1. The absolute error of u in the L^2 -norm for different mesh sizes and $\varepsilon = 1$ or $\varepsilon = 10^{-10}$, using the singular perturbation scheme (P) and the two proposed AP-schemes for a time step of $\tau = 10^{-6}s$ and at instant $t = 10^{-4}$, with $T_m = 1$.

summarized in Table 2 and Figures 4 and 5. Note that the (RK_{AP}) scheme is of second order in time as long as the error due to the time discretization dominates the error induced by the space discretization. The standard (P)-scheme works well and is of first order, as long as ε is close to one. The (E_{AP}) scheme is of first order for all values of the anisotropic parameter. Also note that while the (RK_{AP}) scheme demands twice more computational time than the (E_{AP}) scheme, it gives much better precision. In order to achieve a relative error of the order of 10^{-4} for $\varepsilon = 1$ it suffices to take a time step of $\tau = 0.05$ in the RK-scheme. A comparable accuracy with (E_{AP}) is obtained for a time step 16 times smaller. In the case of $\varepsilon = 10^{-10}$ the ratio is 32.

To conclude, one can remark that the asymptotic-preserving schemes, (E_{AP}) and (RK_{AP}), are uniformly accurate with respect to the perturbation parameter ε . This essential feature can be very useful in situations where the anisotropy is variable in space, *i.e.* the parameter $\varepsilon(x)$ is x -dependent. No mesh-adaptation is any more needed in these cases, a simple Cartesian grid enables accurate results, with no regard to the ε -values.

4.2.2. Initial Gaussian peak.

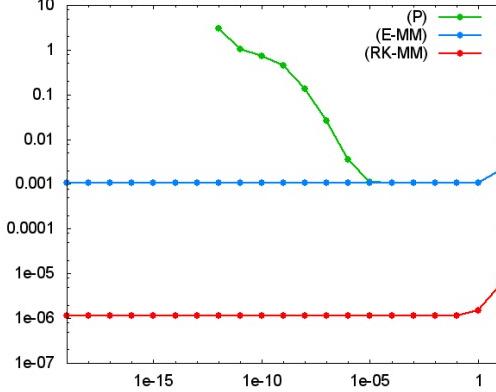


FIGURE 4. Relative L^2 -errors between the exact solution u^ε and the computed solution with the standard scheme (P), the Euler-AP method (E_{AP}) and the DIRK-AP scheme (RK_{AP}) as a function of ε and for $\tau = 0.00625$. The spacial grid is 200×200 .

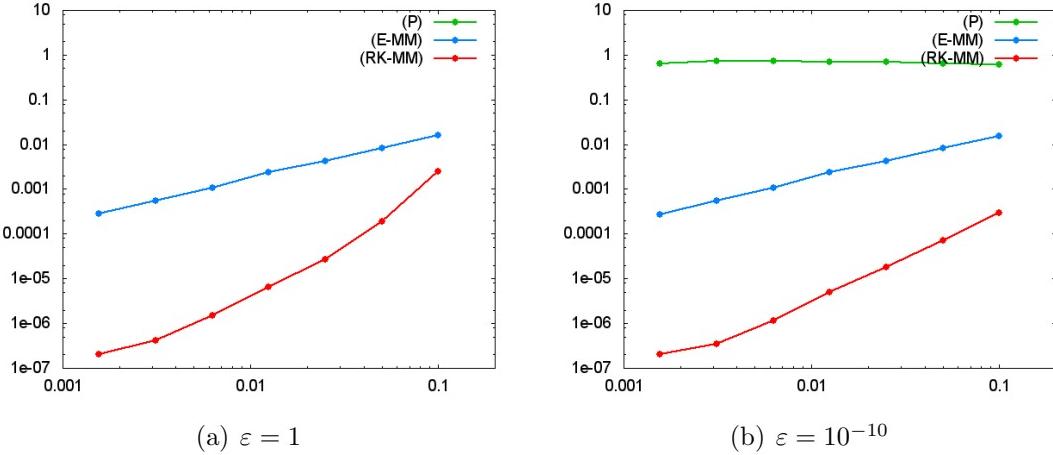


FIGURE 5. Relative L^2 -errors between the exact solution u^ε and the computed solution with the standard scheme (P), the Euler-AP method (E_{AP}) and the DIRK-AP scheme (RK_{AP}) as a function of τ , and for $\varepsilon = 1$ resp. $\varepsilon = 10^{-10}$ and a mesh with 200×200 points. Note that for $\varepsilon = 1$ the P-scheme and the E_{AP} scheme give the same precision.

The second investigated test is the evolution of the following initial Gaussian peak, located in the middle of the computational domain:

$$u(t=0) = \frac{T_m}{2} \left(1 + e^{-50(x-0.5)^2 - 50(y-0.5)^2} \right), \quad (4.53)$$

τ	L^2 -error $\varepsilon = 1$		
	P	E_{AP}	RK_{AP}
0.1	1.57×10^{-2}	1.57×10^{-2}	2.52×10^{-3}
0.05	8.28×10^{-3}	8.28×10^{-3}	1.93×10^{-4}
0.025	4.25×10^{-3}	4.25×10^{-3}	2.62×10^{-5}
0.0125	2.37×10^{-3}	2.37×10^{-3}	6.54×10^{-6}
0.00625	1.08×10^{-3}	1.08×10^{-3}	1.50×10^{-6}
0.003125	5.44×10^{-4}	5.44×10^{-4}	4.08×10^{-7}
0.0015625	2.76×10^{-4}	2.76×10^{-4}	2.07×10^{-7}
τ	L^2 -error $\varepsilon = 10^{-10}$		
	P	E_{AP}	RK_{AP}
0.1	6.14×10^{-1}	1.57×10^{-2}	2.90×10^{-4}
0.05	6.30×10^{-1}	8.22×10^{-3}	7.21×10^{-5}
0.025	6.92×10^{-1}	4.22×10^{-3}	1.80×10^{-5}
0.0125	7.08×10^{-1}	2.36×10^{-3}	4.91×10^{-6}
0.00625	7.26×10^{-1}	1.08×10^{-3}	1.15×10^{-6}
0.003125	7.42×10^{-1}	5.40×10^{-4}	3.43×10^{-7}
0.0015625	6.42×10^{-1}	2.74×10^{-4}	2.05×10^{-7}

TABLE 2. The absolute error of u in the L^2 -norm for different time step using the singular perturbation scheme (P) and two proposed AP-schemes for mesh size 200×200 at time $t = 0.1s$ with $T_m = 1$.

where $T_m = 10^5 K$ is the maximal temperature in the domain and the anisotropy direction is given as in the previous tests. We perform numerical experiments with the choice of $\varepsilon = 1$. We choose the time step $\tau = 0.01$ and perform numerical simulations on a fixed 50×50 grid with the final time set to $15s$. The time step is big compared to the time scale induced by the initial condition. Indeed, after the first iteration of the algorithm the numerical solution immediately falls into the space of functions which are almost constant in the direction of the anisotropy (see Figure 7). The evolution of the numerical solution consists of two phases. The first one, in which the parallel components of the diffusion operator dominate, is characterized by the exponential decay of $\|u_h\|_{L_2(\Omega)}$, $\min(u_h)$ and $\max(u_h)$ (see Figure 6). When u_h reaches some critical value, the parallel part of the diffusion operator becomes smaller than the perpendicular one. The direction of the strong diffusion is now inverted and the numerical solution aligns itself rather with the perpendicular direction. The minimum, maximum as well as the L^2 -norm of u_h continue

to approach zero, but the decay is no longer exponential. The L^2 -norm and the maximal value remain close to each other and almost constant in time. The minimal value of u_h , as well as the boundary-values decrease much faster.

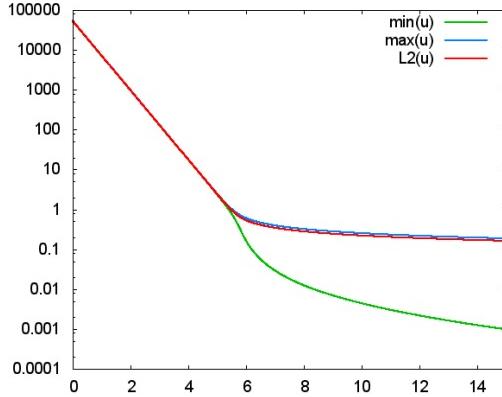


FIGURE 6. $\min(u_h)$, $\max(u_h)$ and $\|u_h\|_{L^2(\Omega)}$ as a function of time for the Gaussian peak experiment, for $T_m = 10^5$ and $\varepsilon = 1$. Time step is $\tau = 0.01s$ and a mesh size of 50×50 .

5. CONCLUSION

The here presented Asymptotic-Preserving scheme proves to be an efficient, general and easy to implement numerical method for solving nonlinear, strongly anisotropic parabolic problems. This kind of problems occur in several important applications, as for example magnetically confined fusion plasmas. The method is based on a reformulation of the problem, initially introduced by the authors in an elliptic framework, and a careful linearization as well as time-discretization of the resulting equation, which does not destroy the AP-properties of the space-discretization. Numerical experiments show clearly the advantages of such an AP-scheme.

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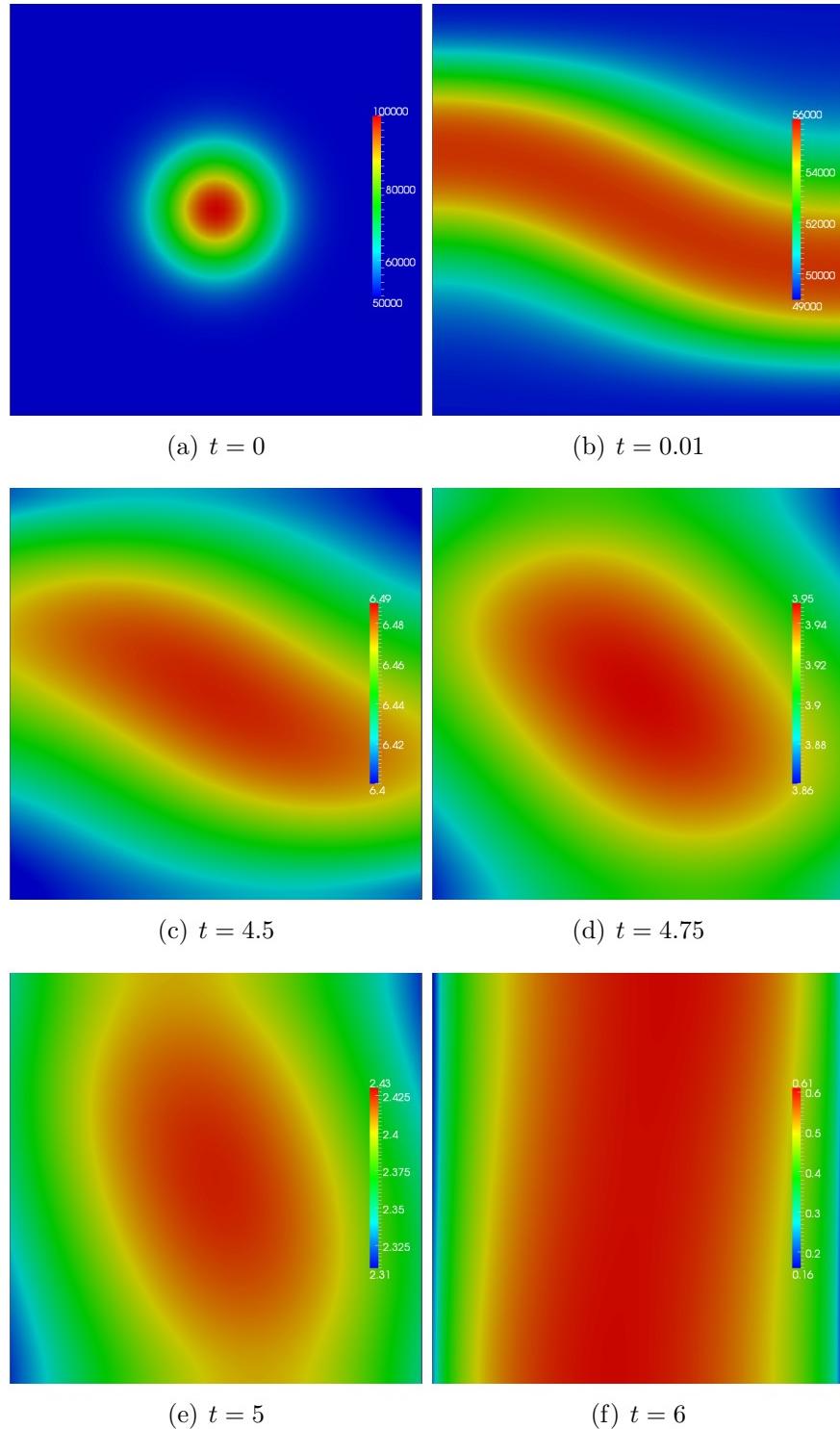


FIGURE 7. Numerical solution at different time steps for the Gaussian peak experiment, for $T_m = 10^5$ and $\varepsilon = 1$. Time step is $\tau = 0.01s$ and a mesh size of 50×50 .

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